Efficient algorithms for differential equation satisfied by Feynman integrals

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with Charles Doran, Andrew Harder, Pierre Lairez, Eric Pichon-Pharabod

and work to appear with Leonardo de la Cruz

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Scattering amplitudes are the fundamental tools for making contact between quantum field theory description of nature and experiments

Comparing particule physics model against datas from accelators

- Post-Minkowskian expansion for Gravitational wave physics
- Various condensed matter and statistical physics systems

Feynman Integrals: parametric representation

Feynman integral are given by projective space integrals

$$I_{\Gamma}(\underline{\nu}, D; \underline{s}, \underline{m}) = \int_{\Delta_n} \frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega - \frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}} \prod_{i=1}^n x_i^{\nu_i - 1} \Omega_0 \qquad \omega = \sum_{i=1}^n \nu_i - \frac{LD}{2}$$

with the volume form on \mathbb{P}^{n-1}

$$\Omega_0 = \sum_{i=1}^n (-1)^{i-1} x^i dx^1 \wedge \cdots \widehat{dx^i} \cdots \wedge dx^n$$

The domain of integration is the positive quadrant

$$\Delta_n := \{x_1 \ge 0, \ldots, x_n \ge 0 | [x_1, \ldots, x_n] \in \mathbb{P}^{n-1}\}$$

Feynman Integrals: parametric representation

The graph polynomial is homogeneous degree L + 1 in \mathbb{P}^{n-1} $\mathcal{F}_{\Gamma}(\underline{x}) = \mathcal{U}_{\Gamma}(\underline{x}) \times \mathcal{L}(\underline{m}^2; \underline{x}) - \mathcal{V}_{\Gamma}(\underline{s}, \underline{x})$

▶ Homogeneous polynomial of degree *L* with $u_{a_1,...,a_n} \in \{0,1\}$

$$\mathcal{U}_{\Gamma}(\underline{x}) := \sum_{\substack{a_1 + \dots + a_n = L\\ \mathbf{0} \le a_i \le 1}} u_{a_1, \dots, a_n} \prod_{i=1}^n x_i^{a_i}$$

the mass hyperplane

$$\mathcal{L}(\underline{m}^2;\underline{x}) := \sum_{n=1}^n m_i^2 x_i$$

Homogeneous polynomial of degree L + 1

$$\mathcal{V}_{\Gamma}(\underline{x}) := \sum_{\substack{a_1 + \dots + a_n = L+1\\ \mathbf{0} < a_i < 1}} S_{a_i, \dots, a_n} \prod_{i=1}^n x_i^{a_i}$$

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The integrand is an algebraic differential form in $H^{n-1}(\mathbb{P}^{n-1}\setminus\mathbb{X}_{\Gamma})$ on the complement of the graph hypersurface

$$\mathbb{X}_{\mathsf{\Gamma}} := \{\mathcal{U}_{\mathsf{\Gamma}}(\underline{x}) imes \mathcal{F}_{\mathsf{\Gamma}}(\underline{x}) = 0, \underline{x} \in \mathbb{P}^{n-1}\}$$

- All the singularities of the Feynman integrals are located on the graph hypersurface
- Generically the graph hypersurface has non-isolated singularities

Feynman integral and periods

The domain of integration Δ_n is not an homology cycle because $\partial \Delta_n \cap \mathbb{X}_{\Gamma} = \{(1, 0, \dots, 0), (0, 1, 0, \dots, 0), \dots, (0, \dots, 0, 1)\}$ we have to look at the relative cohomology

 $H^{\bullet}(\mathbb{P}^{n-1} \setminus X_{\Gamma}; \mathfrak{A}_n \setminus \mathfrak{A}_n \cap \mathbb{X}_{\Gamma})$

The normal crossings divisor $\Pi_n := \{x_1 \cdots x_n = 0\}$ and \mathbb{X}_{Γ} are separated by performing a series of iterated blowups of the complement of the graph hypersurface [Bloch, Esnault, Kreimer]



The Feynman integral *are* period integrals of the relative cohomology after performing the appropriate blow-ups

$$\mathfrak{M}(\underline{s},\underline{m}^2) := H^{\bullet}(\widetilde{\mathbb{P}^{n-1}} \backslash \widetilde{X_F}; \widetilde{\amalg_n} \backslash \widetilde{\amalg_n} \cap \widetilde{X_{\Gamma}})$$

Since the integrand varies with the physical variables $\{S_{\underline{a}^i}, m_1^2, \dots, m_n^2\}$ one needs to study a variation of (mixed) Hodge structure

One can show that the Feynman integral are **holonomic D-finite functions** [Bitoun et al.; Smirnov et al.]

A Feynman integrals satisfies inhomogenous differential equations with respect to any set of variables $\underline{z} \in \{S_{\underline{a}}, m_1^2, \dots, m_n^2\}$

$$\mathscr{L}_{\Gamma}(\underline{z}) I_{\Gamma} = \mathscr{S}_{\Gamma}$$

Generically there is an inhomogeneous term $\mathscr{S}_{\Gamma} \neq 0$ due to the boundary components $\partial \Delta_n$

We want to address the questions

- **1** To what class of functions belong Feynman integrals?
- **2** What is the geometrical algebraic origin of the motive $\mathfrak{M}(\underline{s}, \underline{m}^2)$?
- Derivation of the (D-module of) differential equations ? $\mathscr{L}_{\Gamma}(\underline{z}) l_{\Gamma} = \mathscr{L}_{\Gamma}$

In this talk we focus in the question <a> and present some new methods for deriving such system of differential equation and its underlying (algebraic) geometry

Feynman Integrals differential equations

For a given subset of the physical parameters $\underline{z} := (z_1, \ldots, z_r) \subset \{\underline{s}, \underline{m}^2\}$ we want to derive **minimal order** differential equations

$$\mathscr{L}_{\Gamma}(\underline{s},\underline{m}^{2},\partial_{\underline{z}})\int_{\sigma}\frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega-\frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}}\prod_{i=1}^{n}x_{i}^{\nu_{i}-1}\Omega_{0}=\mathscr{S}_{\sigma,\Gamma}(\underline{z})$$

One way to achieve this is to construct a Gröbner basis of operators $T_{\underline{z}}$ that annihilate the integrand of the Feynman integral

$$T_{\underline{z}}\left(\frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega-\frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}}\prod_{i=1}^{n}x_{i}^{\nu_{i}-1}\Omega_{0}\right)=0$$

such that

$$T_{\underline{z}} = \mathscr{L}_{\Gamma}(\underline{s},\underline{m}^2,\underline{\partial}_{\underline{z}}) + \sum_{i=1}^n \partial_{x_i} Q_i(\underline{s},\underline{m}^2,\underline{\partial}_{\underline{z}};\underline{x},\underline{\partial}_{\underline{x}})$$

Feynman Integrals differential equations

where the finite order differential operator

$$\mathscr{L}_{\Gamma}(\underline{s},\underline{m}^{2},\underline{\partial}_{\underline{z}}) = \sum_{\substack{0 \leq a_{i} \leq o_{i} \\ 1 \leq i \leq r}} p_{a_{1},...,a_{r}}(\underline{s},\underline{m}^{2}) \prod_{i=1}^{r} \left(\frac{d}{dz_{i}}\right)^{a_{i}}$$

$$Q_{i}(\underline{s},\underline{m}^{2},\underline{\partial}_{\underline{z}}) = \sum_{\substack{0 \le a_{i} \le o_{i}^{\prime} \\ 1 \le i \le r}} \sum_{\substack{0 \le b_{i} \le \tilde{\sigma}_{i} \\ 1 \le i \le n}} q_{a_{1},\ldots,a_{r}}(\underline{s},\underline{m}^{2},\underline{x}) \prod_{i=1}^{r} \left(\frac{d}{dz_{i}}\right)^{a_{i}} \prod_{i=1}^{n} \left(\frac{d}{dx_{i}}\right)^{b_{i}}$$

- The orders o_i , o'_i , \tilde{o}_i are positive integers
- ▶ $p_{a_1,...,a_r}(\underline{S},\underline{m}^2)$ polynomials in the kinematic variables

q⁽ⁱ⁾_{a1,...,ar}(<u>s</u>, <u>m</u>², <u>x</u>) rational functions in the kinematic variable and the projective variables <u>x</u>.

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Feynman Integrals differential equations

Integrating over a cycle γ gives

$$0 = \oint_{\gamma} T_{\underline{z}} \Omega_{\Gamma} = \mathscr{L}_{\Gamma}(\underline{s}, \underline{m}, \partial_{\underline{z}}) \oint_{\gamma} \Omega_{\Gamma} + \oint_{\gamma} d\beta_{\Gamma}$$

For a cycle $\partial \gamma = \emptyset$ then $\oint_{\gamma} d\beta_{\Gamma} = 0$ and we get

$$\mathscr{L}_{\Gamma}(\underline{s},\underline{m},\partial_{\underline{z}})\oint_{\gamma}\Omega_{\Gamma}=0$$

For the Feynman integral I_{Γ} we have

$$0 = \int_{\Delta_n} T_{\underline{z}} \Omega_{\Gamma} = \mathscr{L}_{\Gamma}(\underline{s}, \underline{m}, \partial_{\underline{z}}) I_{\Gamma} + \int_{\Delta_n} d\beta_{\Gamma}$$

since $\partial \Delta_n \neq \emptyset$

$$\mathscr{L}_{\Gamma}(\underline{s},\underline{m},\partial_{\underline{z}})I_{\Gamma} = \mathscr{S}_{\Gamma}$$

So we need the telescoper and the certificate

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Feynman integrals Picard–Fuchs

The Rational case

We start with the case of a rational differential form with $D \in 2\mathbb{N}^*$

$$\Omega_{\Gamma} = \frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega - \frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}} \prod_{i=1}^{n} x_{i}^{\nu_{i}-1} \Omega_{0} \qquad \omega = \sum_{i=1}^{n} \nu_{i} - \frac{LD}{2}$$

• We as well assume that all the mass parameters are all vanishing $m_1, \cdots, m_n \neq 0$

• And that $\omega > 0$, i.e. $\sum_{i=1}^{n} \nu_i > LD/2$

So that the integral of Ω_{Γ} on the positive orthan is a convergent integral

The sunset graph

The two-loop sunset graph in D = 2



$$I_{\odot}(p^2,\underline{m}^2) = \int_{\mathbb{R}^3_+} \frac{dx_1 dx_2 dx_3}{\mathcal{F}_{\odot}(\underline{x})}$$

The polar hypersurface of the integral is an elliptic curve $\mathcal{F}_{\odot}(\underline{x}) = 0$

 $\mathcal{F}_{\odot}(\underline{x}) = (x_1x_2 + x_1x_3 + x_2x_3)(m_1^2x_1 + m_2^2x_2 + m_3^2x_3) - p^2x_1x_2x_3$

One can obtain a differential equation annihilating acting on the integral using the Griffiths-Dwork method

Let define the integrand in differential form

$$\Omega_{\odot} = rac{x_1 dx_2 \wedge dx_3 - x_2 dx_1 \wedge dx_3 + x_3 dx_1 \wedge dx_2}{\mathcal{F}_{\odot}(\underline{x})} = rac{\Omega_0}{\mathcal{F}_{\odot}(\underline{x})}$$

consider

$$\frac{\partial\Omega_{\odot}}{\partial p^2} = x_1 x_2 x_3 \frac{\Omega_0}{\mathcal{F}_{\odot}(\underline{x})^2}; \qquad \frac{\partial^2\Omega_{\odot}}{(\partial p^2)^2} = 2(x_1 x_2 x_3)^2 \frac{\Omega_0}{\mathcal{F}_{\odot}(\underline{x})^3}$$

Since we know we have the geometry of an elliptic curve we are looking for a second order differential operator acting on η_{\odot}

$$\mathscr{L}_{\odot}(p^2) = rac{\partial^2}{(\partial p^2)^2} + q_1(p^2, \underline{m}^2) rac{\partial}{\partial p^2} + q_0(p^2, \underline{m}^2)$$

Remark that $(x_1x_2x_3)^2$ lies in the Jacobian ideal for $\mathcal{F}_{\ominus}(\underline{x})$

$$(x_1x_2x_3)^2 = \sum_{i=1}^3 C_i^{(1)}(\underline{x})\partial_{x_i}\mathcal{F}_{\odot}(\underline{x})$$

with $C_i^{(1)}(\underline{x})$ homogeneous of degree 4 in the (x_1, x_2, x_3) variables Following Griffiths one introduces the differential form

$$\beta_{1} = \frac{(x_{2}C_{3}^{(1)}(\underline{x}) - x_{3}C_{2}^{(1)}(\underline{x}))dx_{1}}{\mathcal{F}_{\odot}(\underline{x})^{2}} + \frac{(x_{1}C_{3}^{(1)}(\underline{x}) - x_{3}C_{1}^{(1)}(\underline{x}))dx_{2}}{\mathcal{F}_{\odot}(\underline{x})^{2}} + \frac{(x_{1}C_{2}^{(1)}(\underline{x}) - x_{2}C_{1}^{(1)}(\underline{x}))dx_{3}}{\mathcal{F}_{\odot}(\underline{x})^{2}}$$

such that

$$d\beta_1 = 2 \frac{\sum_{i=1}^3 C_i^{(1)}(\underline{x}) \partial_{x_i} \mathcal{F}_{\odot}(\underline{x}) \Omega_0}{\mathcal{F}_{\odot}(\underline{x})^3} - \frac{\sum_{i=1}^3 \partial_{x_i} C_i^{(1)}(\underline{x}) \Omega_0}{\mathcal{F}_{\odot}(\underline{x})^2}$$

$$\mathscr{L}_{\odot}(p^2)\Omega_{\odot} = rac{q_1(p^2,\underline{m}^2)x_1x_2x_3+\sum_{i=1}^3\partial_{x_i}C_i^{(1)}(\underline{x})}{\mathcal{F}_{\odot}(\underline{x})^2}\Omega_0 + deta_1$$

We can again reduce this second order pole using that there exist a polynomial $q_1(p^2, \underline{m}^2)$ such that

$$q_1(p^2,\underline{m}^2)x_1x_2x_3+\sum_{i=1}^3\partial_{x_i}C_i^{(1)}(\underline{x})=\sum_{i=1}^3C_i^{(2)}\partial_{x_i}\mathcal{F}_{\odot}(\underline{x})$$

with $C_i^{(2)}$ of degree 1. One introduces the 1-form β_2

$$\beta_2 = \sum_{i=1}^{3} \epsilon^{ijk} \frac{x_j C_k^{(2)}(\underline{x}) dx_i}{\mathcal{F}_{\odot}(\underline{x})}$$

such that

$$d\beta_{2} = \frac{\sum_{i=1}^{3} C_{i}^{(2)}(\underline{x}) \partial_{x_{i}} \mathcal{F}_{\ominus}(\underline{x})}{\mathcal{F}_{\ominus}(\underline{x})^{2}} - \frac{\sum_{i=1}^{3} \partial_{x_{i}} C_{i}^{(2)}(\underline{x}) \Omega_{0}}{\mathcal{F}_{\ominus}(\underline{x})}$$

We have achieved that

$$\left(\frac{\partial^2}{(\partial p^2)^2} + q_1(p^2, \underline{m}^2)\frac{\partial}{\partial p^2} + \sum_{i=1}^3 \partial_{x_i} C_i^{(2)}(\underline{x})\right) \eta_{\ominus} = d(\beta_1 + \beta_2)$$

because the $C_i^{(2)}(\underline{x})$ are of degree 1 in (x_1, x_2, x_3) then $q_0(p^2, \underline{m}^2) = \partial_{x_i} C_i^{(2)}(\underline{x})$ only depends on p^2, \underline{m}^2

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We then conclude that the minimal operator acting on the sunset integral is the Picard-Fuchs operator

$$\mathscr{L}_{p^2} = \frac{\partial^2}{(\partial p^2)^2} + q_1(p^2, \underline{m}^2) \frac{\partial}{\partial p^2} + q_0(p^2, \underline{m}^2)$$

which acts on the integrals as

$$\mathscr{L}_{p^2} I_{\odot}(p^2) = \int_{x_i \ge 0} \mathscr{L}_{p^2} \frac{\Omega_0}{\mathcal{F}_{\odot}(\underline{x})} = \int_{x_i \ge 0} d(\beta_1 + \beta_2) \neq 0$$

We have constructed by the telescoper $T_{p^2} = \mathscr{L}_{\odot}(p^2)$ and the certificate $C_{\odot} = d(\beta_1 + \beta_2)$

The differential operator \mathscr{L}_{p^2} is the Picard–Fuchs operator of the elliptic curve defined by the graph polynomial $\mathcal{F}_{\ominus}(x_1, x_2, x_3) = 0$

If one considers the family of elliptic curve E

 $y^2 = 4x^3 - g_2(t)x - g_3(t);$ $\Delta(t) = g_2(t)^3 - 27g_3(t)^2$

the periods satisfy the differential system of equations

$$\frac{d}{dt} \begin{pmatrix} \int_{\gamma} \frac{dx}{y} \\ \int_{\gamma} \frac{xdx}{y} \end{pmatrix} = \begin{pmatrix} -\frac{1}{12} \frac{d}{dt} \log \Delta(t) & \frac{3\delta(t)}{2\Delta(t)} \\ -\frac{g_2(t)\delta(t)}{8\Delta(t)} & \frac{1}{12} \frac{d}{dt} \log \Delta(t) \end{pmatrix} \begin{pmatrix} \int_{\gamma} \frac{dx}{y} \\ \int_{\gamma} \frac{xdx}{y} \end{pmatrix}$$

with $\delta(t) = 3g_3(t) \frac{d}{dt}g_2(t) - 2g_2(t) \frac{d}{dt}g_3(t)$ The Picard–Fuchs operator acting on the period integral $\int_{\gamma} dx/y$ is

$$\begin{aligned} \mathscr{L}_{\text{ell}} &= 144\Delta(t)^{2}\delta(t)\frac{d^{2}}{dt^{2}} + 144\Delta(t)\left(\delta(t)\frac{d\Delta(t)}{dt} - \Delta(t)\frac{d\delta(t)}{dt}\right)\frac{d}{dt} \\ &+ 27g_{2}(t)\delta(t)^{3} + 12\frac{d^{2}\Delta(t)}{dt^{2}}\delta(t)\Delta(t) - \left(\frac{d\Delta(t)}{dt}\right)^{2}\delta(t) - 12\frac{d\delta(t)}{dt}\Delta(t)\frac{d\Delta(t)}{dt}. \end{aligned}$$

This matches the differential operator derived using the Griffiths–Dwork method

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In general the graph hypersurface does not have isolated singularities (which is the generic case) therefore the "naïve" implementation of the Griffiths-Dwork algorithm does not work

One could use the implementation of Doron Zeilberger (1990) creative telescoping algorithm by F. Chyzak or C. Koutschan but the algorithm takes a very long time for graph with many edges

$$\Omega_{\Gamma} = \frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega - \frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}} \prod_{i=1}^{n} x_{i}^{\nu_{i} - 1} \Omega_{0}$$

We used **period** by Pierre Lairez of an extended Griffiths-Dwork algorithm that handles singular hypersurfaces

Extended Griffiths-Dwork: syzygies

In this example we saw the pole reduction

$$\frac{\partial^2 \eta_{\odot}}{(\partial \rho^2)^2} = \frac{2(x_1 x_2 x_3)^2}{\mathcal{F}_{\odot}(\underline{x})^3} \Omega_0 = \frac{\sum_{i=1}^3 \partial_{x_i} C_i}{\mathcal{F}_{\odot}(\underline{x})^2} \Omega_0 + d\beta_1$$

For singular hypersurface $X_{\Gamma} \subset \mathbb{P}^{n-1}$ the Jacobian reduction may not be enough to reduce the pole order when $k \geq n$ Other reduction rules come from the *syzygies* of the derivatives $\frac{\partial \mathcal{F}_{\Gamma}}{\partial x_i}$, i.e.

tuples (B_1, \ldots, B_n) be homogeneous of degree $k \deg \mathcal{F}_{\Gamma} - n + 1$ such that $\sum_i B_i \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_i} = 0$ such that $(\xi_i = (-1)^{i-1} dx_1 \cdots \widehat{dx_i} \cdots dx_n)$

$$\frac{\sum_{i} \frac{\partial B_{i}}{\partial x_{i}}}{\mathcal{F}_{\Gamma}^{k}} \Omega_{0} = d\left(\sum_{i} \frac{B_{i}}{\mathcal{F}_{\Gamma}^{k}} \xi_{i}\right) \Longrightarrow \int_{\gamma} \frac{\sum_{i} \frac{\partial B_{i}}{\partial x_{i}}}{\mathcal{F}_{\Gamma}^{k}} \Omega_{0} = 0.$$

In singular cases, these relations are missed by the Griffiths–Dwork reduction, we need the extended Griffiths–Dwork reduction implemented by [Lairez]

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Feynman integrals Picard–Fuchs

Extended Griffiths-Dwork : syzygies

Given a form $\Omega = \frac{A}{\mathcal{F}_{r}^{k}} d\underline{x}$ deg $A = k \deg \mathcal{F}_{\Gamma} - n$

Compute a basis of the space S_k of all syzygies of degree k deg F_Γ − n + 1 quotiented by the space of trivial syzygies

$$D_{ij} = -D_{ji}, \qquad B_i = \sum_{j=1}^n D_{ij} \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_j} \Longrightarrow \sum_i B_i \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_i} = 0$$

are irrelevant because already used by the Griffiths-Dwork reduction

Extended Griffiths-Dwork : syzygies

Given a form $\Omega = \frac{A}{\mathcal{F}_{\Gamma}^{k}} d\underline{x}$ deg $A = k \deg \mathcal{F}_{\Gamma} - n$ Compute a normal form R of A modulo the Jacobian ideal plus the space $dV = \left\{ \sum_{i} \frac{\partial B_{i}}{\partial x_{i}} \mid \underline{B} \in V \right\}$, that is for some polynomials B_{i} and C_{i} $A = R + \sum_{i} \frac{\partial B_{i}}{\partial x_{i}} + \underbrace{C_{1} \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_{1}} + \dots + C_{n} \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_{n}}}_{\in \text{Jacobian ideal}}$

Extended Griffiths-Dwork : syzygies

Given a form $\Omega = \frac{A}{\mathcal{F}_{r}^{k}} d\underline{x}$ deg $A = k \deg \mathcal{F}_{\Gamma} - n$

On the second second

$$(k-1)\frac{A}{\mathcal{F}_{\Gamma}^{k}}d\underline{x} = \frac{\sum_{i}\frac{\partial C_{i}}{\partial x_{i}}}{\mathcal{F}_{\Gamma}^{k-1}}d\underline{x} - d\left(\sum_{i}\frac{B_{i}}{\mathcal{F}_{\Gamma}^{k}}\xi_{i} + \sum_{i}\frac{C_{i}}{\mathcal{F}_{\Gamma}^{k-1}}\xi_{i}\right).$$

Then

$$\int_{\gamma} \frac{A(\underline{x})}{\mathcal{F}_{\Gamma}(\underline{x})^{k}} d\underline{x} = -\frac{1}{k-1} \int_{\gamma} \frac{\sum_{i} \frac{\partial C_{i}}{\partial x_{i}}}{\mathcal{F}_{\Gamma}^{k-1}} d\underline{x},$$

The extended Griffiths–Dwork reduction presented above is not always enough and may need further extensions, i.e. syzygies of syzygies. There is a hierarchy of extensions which eventually collapse to the strongest possible reduction.

However, for all the computations presented here, we only needed the first extension.

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In the construction we will only consider the case where $\beta(\underline{x}, t)$ is holomorphic on $\mathbb{P}^{n-1} \setminus X_{\Gamma}$, that is is β_{Γ} does not have poles that are not present in Ω_{Γ} .

Consider the rational function $F(x_1, x_2)$

$$\frac{ax_1 + bx_2 + c}{\left(\alpha x_1^2 + \beta x_2^2 + \gamma x_1 x_2 + \delta x_1 + \eta x_2 + \zeta\right)^2} = \partial_{x_1} \frac{N_1(x_1, x_2)}{D_1(x_1, x_2)} + \partial_{x_2} \frac{N_2(x_1, x_2)}{D_2(x_1, x_2)}$$

where $a, b, c, \alpha, \beta, \gamma, \delta, \eta, \eta$ are constants and polynomials $N_i(x_1, x_2)$ and $D_i(x_1, x_2)$ with i = 1, 2. The denominators have poles at $x_2^0 = (a\delta - 2\alpha c)/(2\alpha b - a\gamma)$ which is not a pole of the left-hand-side. This means one can find a cycle γ passing by x_2^0 such that the integral of $\int_{\gamma} F(x_1, x_2)$ is finite and non-vanishing.

Minimality of the Picard–Fuchs operator

This dimension dim $(V_{\Gamma}) = (-1)^{n+1}\chi((\mathbb{C}^*)^n \setminus \mathbb{V}(\mathcal{U}_{\Gamma}) \cup \mathbb{V}(\mathcal{F}_{\Gamma}))$ gives an upper bound on the order of the minimal order differential operator acting on the Feynman integral. [Bitour et al.]

The extended Griffiths–Dwork algorithm leads to a minimal order differential operator

 $\mathscr{L}_{\Gamma}\Omega_{\Gamma} = d\beta_{\Gamma}$

annihilating (in cohomology) the Feynman integral differential form Ω_{Γ} with the condition that the certificate β_{Γ} is an holomorphic form on $\mathbb{P}^{n-1} \setminus Z_{\Gamma}$. \mathscr{L}_{Γ} is the minimal differential order differential operator satisfying this condition.

Using the algorithm by [Chyzak, Goyer, Mezzarobba] we test the irreducibility of the the Picard–Fuchs operator and factorize when it is reducible.

Sunset graph Picard–Fuchs operator

$$\begin{array}{c}
 & \Omega_n^{\odot}(t,\underline{m}^2) := \frac{\Omega_0}{\mathcal{F}_n^{\odot}(t,\underline{m}^2;\underline{x})} \in H^{n-1}(\mathbb{P}^{n-1} \setminus X_{\odot}) \\
 & \mathcal{F}_n^{\odot}(t,\underline{m}^2;\underline{x}) := x_1 \cdots x_n \left(\left(\sum_{i=1}^n \frac{1}{x_i} \right) \left(\sum_{j=1}^n m_j^2 x_j \right) - t \right)
\end{array}$$

The graph hypersurface $\mathcal{F}_n^{\ominus}(t, \underline{m}^2; x) = 0$ defines a Calabi-Yau manifold of dimension n-1

For generic physical parameters configurations we find a minimal order Picard–Fuchs operator [Lairez, Vanhove]

$$\mathscr{L}_t = \sum_{r=0}^{o_n} q_r(t,\underline{m}^2) \left(\frac{d}{dt}\right)^r \qquad o_n = 2^n - \binom{n+1}{\lfloor \frac{n+1}{2} \rfloor}; n \ge 2.$$

The three-loop sunset graph



At three-loop we have a K3 surface of 19 < Pic < 16 depending on the mass $m_4=m_1$ For generic mass parameters the

Picard-Fuchs operators

$$\mathscr{L}_6 = \sum_{r=0}^6 q_r(s) \left(rac{d}{dp^2}
ight)^r$$

is order 6 and degree 25

$$egin{aligned} q_6(m{
ho}^2) &= \widetilde{q}_6(m{
ho}^2) imes \ &\prod_{\epsilon_i=\pm 1} (m{
ho}^2 - (\epsilon_1 m_1 + \epsilon_2 m_2 + \epsilon_3 m_3 + \epsilon_4 m_4)^2) \end{aligned}$$

singularities

The four-loop sunset graph



The geometry is the one of the Calabi-Yau threefold. For generic kinematics we have an order 12 degree 121 operator

$${\cal L}_t^{[1^5]} = \sum_{r=0}^{12} q_r^{[1^5]}(t, {\underline m}^2) \left({d \over dt}
ight)^r \, .$$

The degree of the apparent singularities is a polynomial of degree 98

- The six mass configuration m₁ ≠ m₂ ≠ m₃ ≠ m₄ ≠ m₅ ≠ m₆ denote [1⁶]: the Picard–Fuchs operator of order 29 and degree of the polynomial q₂₉(t) is 521.
- ▶ The seven mass configuration $m_1 \neq m_2 \neq m_3 \neq m_4 \neq m_5 \neq m_6 \neq m_7$: the Picard–Fuchs operator of order 58 with a degree 2273
- Results compatible with a CY n 1-fold

Results obtained using Pierre Lairez period

The rational differential form in \mathbb{P}^5

$$\Omega(t) = \frac{\Omega_0^{(6)}}{\left(\mathcal{U}_6(\underline{x})\mathcal{L}_6(\underline{m}^2,\underline{x}) - t\mathcal{V}(\underline{s},\underline{x})\right)^2}$$

 $\underline{P_3} \ \mathcal{U}_6(\underline{x}) = (x_1 + x_2)(x_3 + x_4) + (x_1 + x_2)(x_5 + x_6) \\ + (x_3 + x_4)(x_5 + x_6)$

 $\mathcal{V}(\underline{s},\underline{x}) = \sum_{1 \leq i,j,k \leq 6} C_{ijk} y_i y_j y_k$ with linear changes $(x_{2i-1}, x_{2i}) \rightarrow (y_{2i-1}, y_{2i})$ and i = 1, 2, 3 C_{ijk} symmetric traceless i.e. $C_{iij} = 0$ The algorithm gives an irreducible Picard–Fuchs operator of order 11 with an head polynomial of degree up to 215. [Lairez, Vanhove]



The rational differential form in \mathbb{P}^5

 p_2 Dл p_5

 $\Omega(t) = \frac{\Omega_0^{(6)}}{\left(\mathcal{U}_6(\underline{x})\mathcal{L}_6(\underline{m}^2,\underline{x}) - t\mathcal{V}(\underline{s},\underline{x})\right)^2}$

The motive associated to this graph

Theorem (Tardigrade motive)

Let $X_{(2,2,2);D}$ be the tardigrade hypersurface for generic mass and momentum parameters and $D \ge 2$. Then there is a quartic K3 surface with six A₁ singularities so that $\operatorname{Gr}_4^W \operatorname{H}^4(X_{(2,2,2);D};\mathbb{Q})$ is isomorphic to $\operatorname{H}^2(S;\mathbb{Q})(-1)$ for a K3 surface S up to mixed Tate factors.

Determined in [Doran, Harder, Pichon-Pharabod, Vanhove]



The rational differential form in \mathbb{P}^5

 $\Omega(t) = \frac{\Omega_0^{(6)}}{\left(\mathcal{U}_6(\underline{x})\mathscr{L}_6(\underline{m}^2, \underline{x}) - t\mathcal{V}(\underline{s}, \underline{x})\right)^2}$

The singularities of the graph polynomials are all of type A_1 and one can apply Eric Pichon-Pharabod program lefschetz-family to (numerically) determine the transcendental lattice and confirm that we have a K3 of Picard Rank 11

The non-rational case

We now consider the non-rational case $D \in \mathbb{R}$

$$\Omega_{\Gamma} = \left(\frac{\mathcal{U}_{\Gamma}(\underline{x})}{\mathcal{F}_{\Gamma}(\underline{x})}\right)^{\sum_{i}\nu_{i}} \left(\frac{\mathcal{U}_{\Gamma}(\underline{x})^{L+1}}{\mathcal{F}_{\Gamma}(\underline{x})^{L}}\right)^{D} \prod_{i=1}^{n} x_{i}^{\nu_{i}-1} \Omega_{0}$$

- ▶ We relax all assumption on the mass parameters who can all vanish $m_1, \cdots, m_n \in \mathbb{R}$
- We have degree 0 rational form

$$R(\underline{x}) := rac{\mathcal{U}_{\Gamma}(\underline{x})^{L+1}}{\mathcal{F}_{\Gamma}(\underline{x})^{L}}$$

As a function of the powers of the propagators $\underline{\nu}$ and the dimension D the integral has singularities located on hyperplane defined by $\sum_{i=1}^{n} a_i \nu_i + a_0 D = 0 \text{ with } (a_0, a_1, \dots, a_n) \in \mathbb{Z}^{n+1}$

One can perform a Laurent expansion near say $D_c = 4$ dimensions

$$I_{\Gamma}(\underline{s},\underline{m}^{2};\underline{\nu},D) = \sum_{r\geq -2L} (D-D_{c})^{r} I_{\Gamma}^{(r)}(\underline{s},\underline{m}^{2};\underline{\nu})$$

where $I_{\Gamma}^{(r)}(\underline{s}, \underline{m}^2; \underline{\nu})$ are convergent integrals.

Eugene R. Speer Generalized Feynman Amplitudes Princeton University Press, (1969)

The sunset graph

The two-loop sunset graph in $D = 2 - 2\epsilon$ with $\epsilon \in \mathbb{R}$



$$I_{\odot}(p^2,\underline{m}^2) = \int_{\mathbb{R}^3_+} \left(\frac{\mathcal{U}^3_{\odot}(\underline{x})}{\mathcal{F}^2_{\odot}(\underline{x})}\right)^{\epsilon} \frac{dx_1 dx_2 dx_3}{\mathcal{F}_{\odot}(\underline{x})}$$

with

$$\mathcal{U}_{\odot}(\underline{x}) = x_1 x_2 + x_1 x_3 + x_2 x_3$$

The polar hypersurface of the integral is still the elliptic curve $\mathcal{F}_{\odot}(\underline{x}) = 0$

 $\mathcal{F}_{\odot}(\underline{x}) = (x_1x_2 + x_1x_3 + x_2x_3)(m_1^2x_1 + m_2^2x_2 + m_3^2x_3) - p^2x_1x_2x_3$

We consider differentiation with respect to a single physical parameter $z \in \{\vec{m}, \vec{s}\}$

We consider the derivative

$$\left(\frac{d}{dz}\right)^{a}\Omega_{\Gamma}^{\epsilon} = \frac{\Gamma(1+a+\epsilon)}{\Gamma(1+2\epsilon)}\frac{(x_{1}x_{2}x_{3})^{a}}{\mathcal{F}_{\odot}^{a+1}}\left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon}\,\Omega_{n}^{(0)}$$

we reduce the numerator in the Jacobian ideal of \mathcal{F}_{\ominus}

$$\frac{\Gamma(1+a+\epsilon)}{\Gamma(1+2\epsilon)} (x_1 x_2 x_3)^a = \vec{C}_{(a)} \cdot \vec{\nabla} \mathcal{F}_{\odot}.$$

Integration by part gives

$$\left(\frac{d}{dz}\right)^{a}\Omega_{\Gamma}^{\epsilon} = \frac{\vec{\nabla}\cdot\vec{\mathcal{C}}_{(a)}}{a\mathcal{F}_{\odot}^{a}} \left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon} \Omega_{0} + \epsilon \frac{\vec{\mathcal{C}}_{(a)}\cdot\vec{\nabla}\log(\mathcal{U}_{\odot}^{3}/\mathcal{F}_{\odot}^{2})}{a\mathcal{F}_{\odot}^{a}} \left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon} \Omega_{0} + d\beta_{(a)}$$

or equivalently

$$\left(\frac{d}{dz}\right)^{a}\Omega_{\Gamma}^{\epsilon} = \frac{\vec{\nabla}\cdot\vec{C}_{(a)}}{(a+2\epsilon)\mathcal{F}_{\odot}^{a}} \left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon} \Omega_{0} + 3\epsilon \frac{\vec{C}_{(a)}\cdot\vec{\nabla}\log(\mathcal{U}_{\odot})}{(a+2\epsilon)\mathcal{F}_{\odot}^{a}} \left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon} \Omega_{0} + d\beta_{(a)}$$

We ask that

$$\vec{C}_{(a)} \cdot \vec{\nabla} \mathcal{U}_{\odot} = c_{(a)}(\underline{x})\mathcal{U}_{\odot}.$$

Solving the system

$$\frac{\Gamma(1+a+\epsilon)}{\Gamma(1+2\epsilon)} (x_1 x_2 x_3)^a = \vec{C}_{(a)} \cdot \vec{\nabla} \mathcal{F}_{\ominus}, \\ \vec{C}_{(a)} \cdot \vec{\nabla} \mathcal{U}_{\ominus} = c_{(a)}(\underline{x}) \mathcal{U}_{\ominus}.$$

Gives the pole reduction

$$\left(\frac{d}{dz}\right)^{a}\Omega_{\Gamma}^{\epsilon}=\frac{\vec{\nabla}\cdot\vec{\mathcal{C}}_{(a)}+3\epsilon c_{(a)}(\underline{x})}{(a+2\epsilon)\mathcal{F}_{\odot}^{a}}\left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon}\Omega_{0}+d\beta_{(a)}.$$

- This tells us how to modify the Griffiths-Dwork pole reduction and deduce the e deformed differential equation.
- This allows to treat case that are divergence for
 e = 0 which was not
 possible with the previous algorithm

Work in progress [de la Cruz, Vanhove]

Pierre Vanhove (IPhT)

For the all equal mass case the algorithm gives (up to 20 loops) For the all equal mass case $m_1 = \cdots = m_{l+1} = 1$ we find the sunset Feynman integral satisfies the differential equation

$$\mathscr{L}^{(l),\epsilon}_{\odot}$$
 $l_{\odot}(\{1,\ldots,1\},t,\epsilon) = -(l+1)!rac{\Gamma(1+\epsilon)^l}{\Gamma(1+l\epsilon)}$

with

$$\mathscr{L}_{\odot}^{(l),\epsilon} = \mathscr{L}_{\odot}^{(l),l} + \epsilon \mathscr{L}_{\odot}^{(l),l-1} + \dots + \epsilon^{l} \mathscr{L}_{\odot}^{(l),0}$$

where the differential operator is $\mathscr{L}_{\odot}^{(l),r}$ is of order r.

$$\mathscr{L}_{\odot}^{(2),\epsilon} = \mathscr{L}_{1}^{(1)}\mathscr{L}_{1}^{(2)}\mathscr{L}_{\odot}^{3-\textit{mass}} + \epsilon\mathscr{L}_{4}^{(3)} + \epsilon^{2}\mathscr{L}_{3}^{(4)} + \epsilon^{3}\mathscr{L}_{2}^{(5)} + \epsilon^{4}\mathscr{L}_{1}^{(6)} + \epsilon^{5}\mathscr{L}_{0}^{(7)}$$

where $\mathscr{L}_m^{(r)}$ are irreducible differential operator of order m and $\mathscr{L}_{\ominus}^{3-mass}$ is the differential operator for the three-mass two-loop sunset integral in two dimensions.

Its actions on the Feynman integral is given by

$$\mathscr{L}_{\odot}^{(2),\epsilon} I(\underline{m},t;\epsilon) = \mathscr{S}(\vec{m},t;\epsilon)$$

with the source term

$$\mathscr{S}(\vec{m},t;\epsilon) = \frac{c_{23}(t,\epsilon)\Gamma(\epsilon+1)^2}{(m_2m_3)^{2\epsilon}\Gamma(1+2\epsilon)} + \frac{c_{13}(t,\epsilon)\Gamma(\epsilon+1)^2}{(m_1m_3)^{2\epsilon}\Gamma(1+2\epsilon)} + \frac{c_{12}(t,\epsilon)\Gamma(\epsilon+1)^2}{(m_1m_2)^{2\epsilon}\Gamma(1+2\epsilon)}$$

Two loop Sunset: different masses II

The ϵ deformed operator has for highest order term

$$\begin{aligned} \mathcal{C}_{\ominus}^{(2),\epsilon} \Big|_{(d/dt)^4} &= t^3 \prod_{i=1}^4 (t - \mu_i^2) \\ &\times \left(-(2\epsilon + 5) t^2 - 2 \left(m_1^2 + m_2^2 + m_3^2 \right) (1 + 2\epsilon) t + (7 + 6\epsilon) \prod_{i=1}^4 \mu_i \right) \end{aligned}$$

where

 $\mu_i = \{m_1 + m_2 + m_3, -m_1 + m_2 + m_3, m_1 - m_2 + m_3, m_1 + m_2 - m_3\}$ are the thresholds.

- The ϵ deformation is only affecting the apparent singularities
- The non-apparent singularities are still the roots of the discriminant of the sunset elliptic curve
- The order 4 operator is irreducible

☆ We have put forward a new approach for deriving the differential equation for Feynman integrals

- ☆ We can derive the differential equations in general dimension by extending the Griffiths-Dwork reduction
- $\ref{eq: the see}$ We see how the twist ϵ -factor affects only the apparent singularities
- For graphs with many edges the reduction takes a long time we have been using the FiniteFlow program to speed up the computation but still improvements are needed

We have a seminar on these mathematical aspects of Feynman integral run by Francis Brown, Erik Panzer, Federico Zerbini and myself at the address https://www.ihes.fr/~vanhove/motivefeynman.html